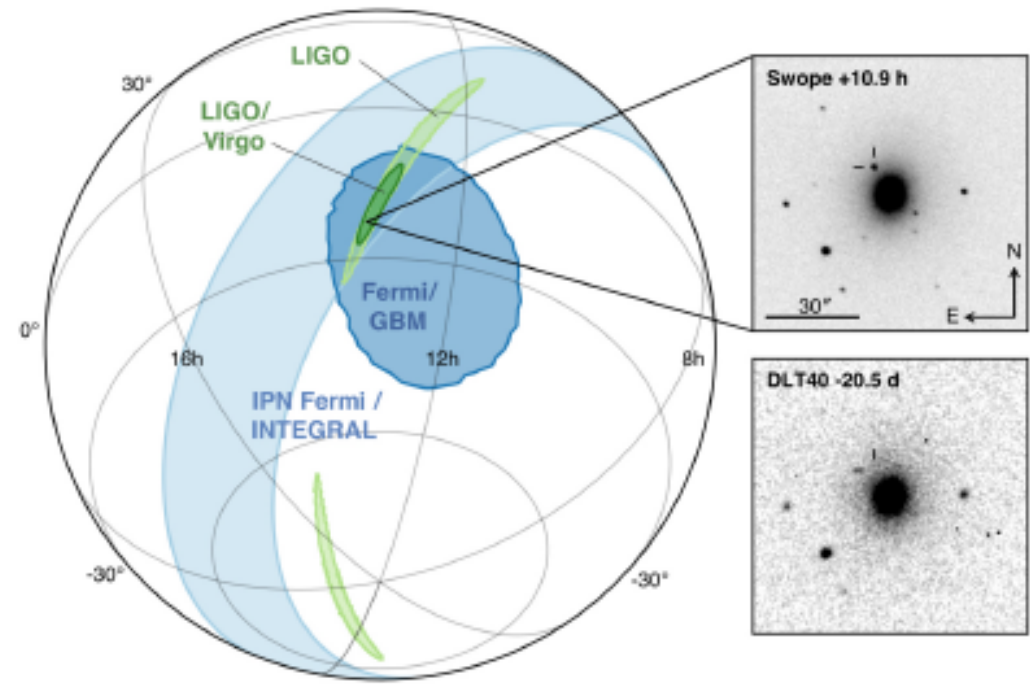
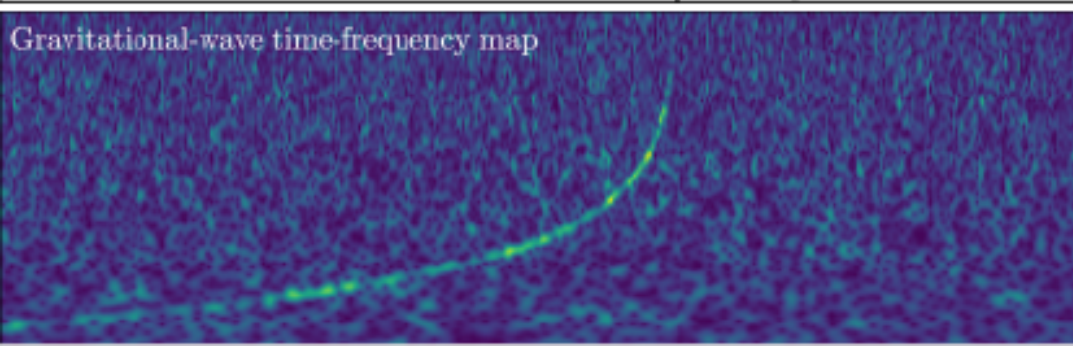
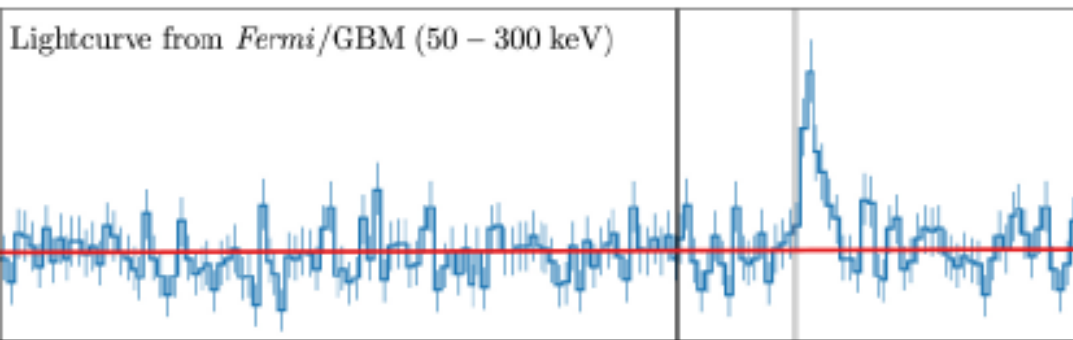
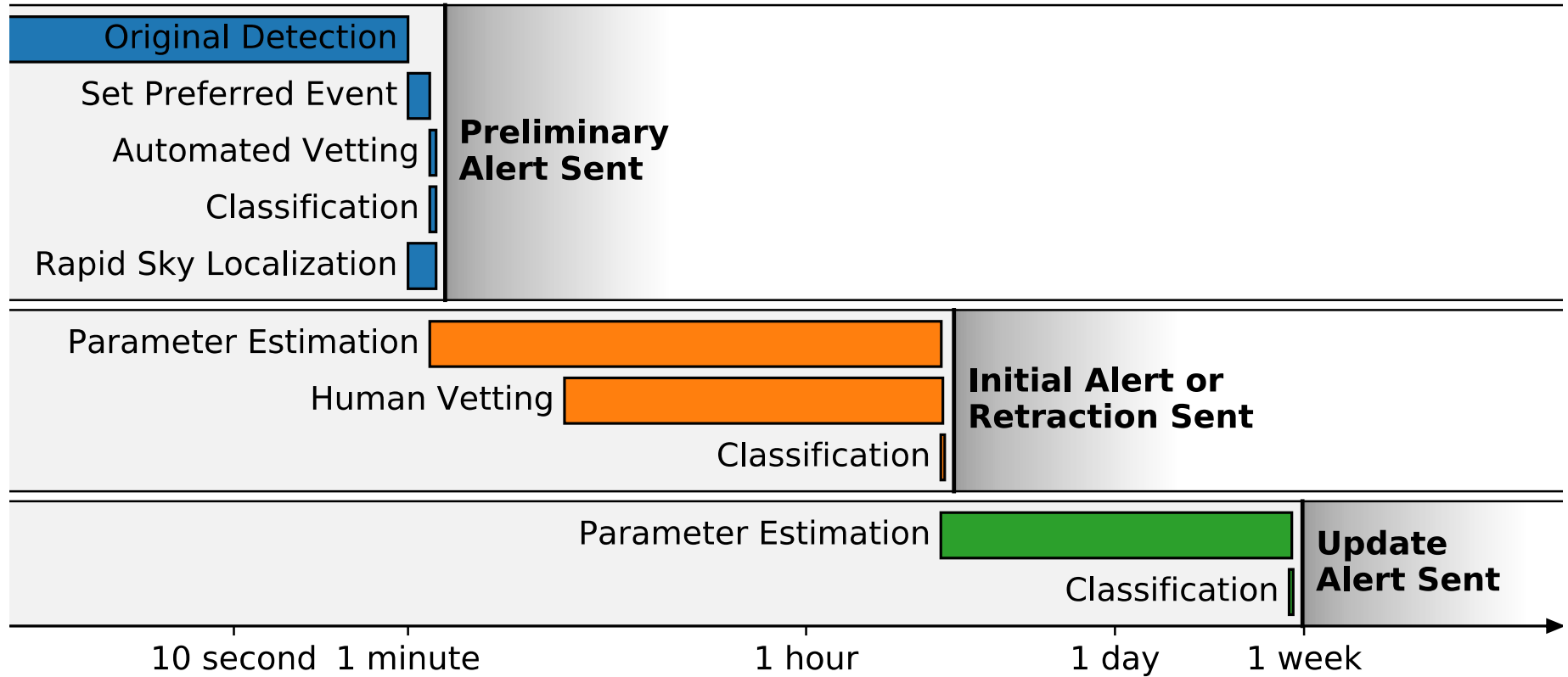


# Open LIGO-Virgo alerts



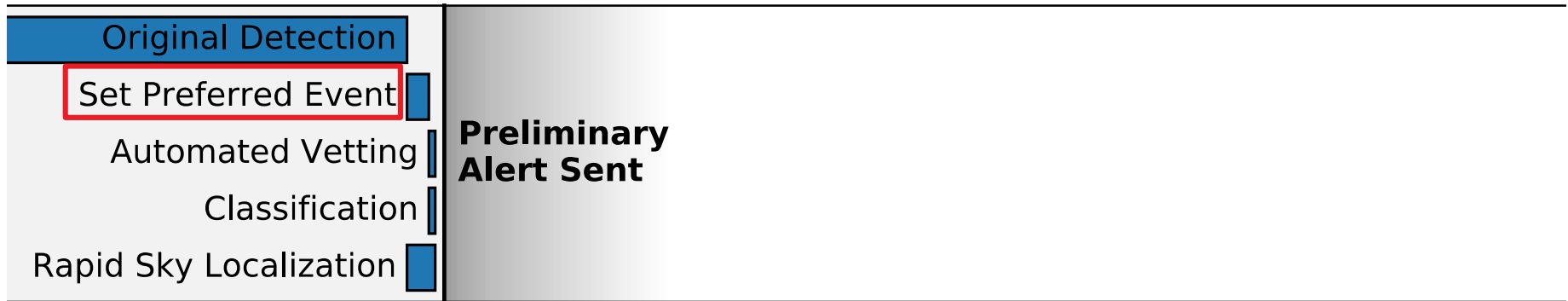
# Timeline

Time since gravitational-wave signal



# Timeline

Time since gravitational-wave signal

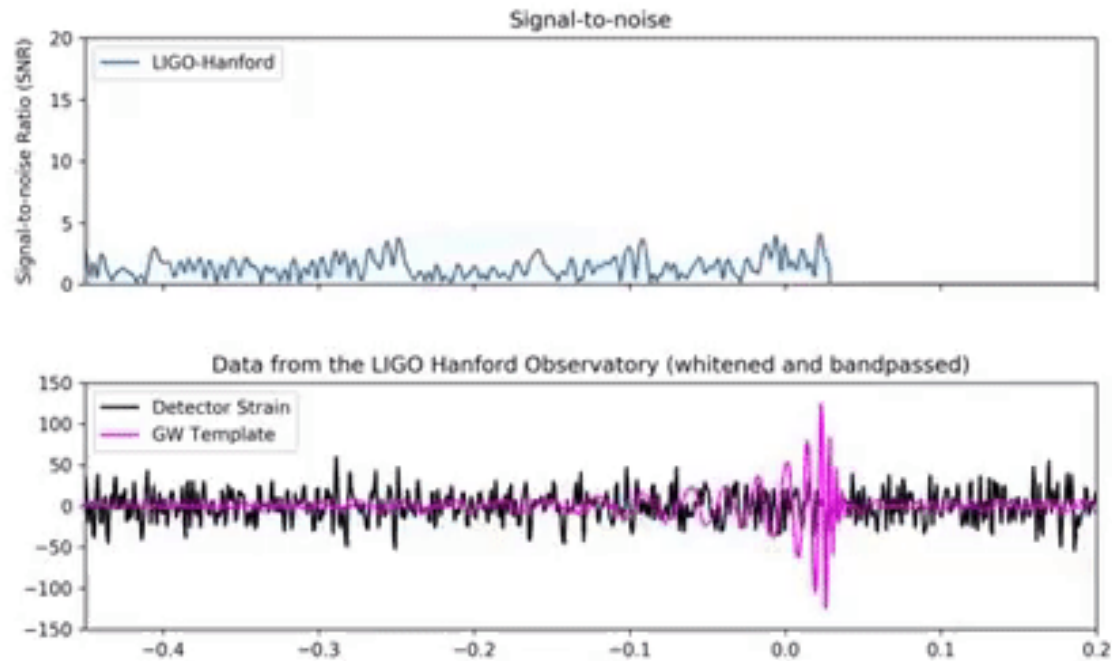


Several pipelines, each can make several detections

Based on a given criteria, one is chosen as being the **Preferred Event** (for the public alert)

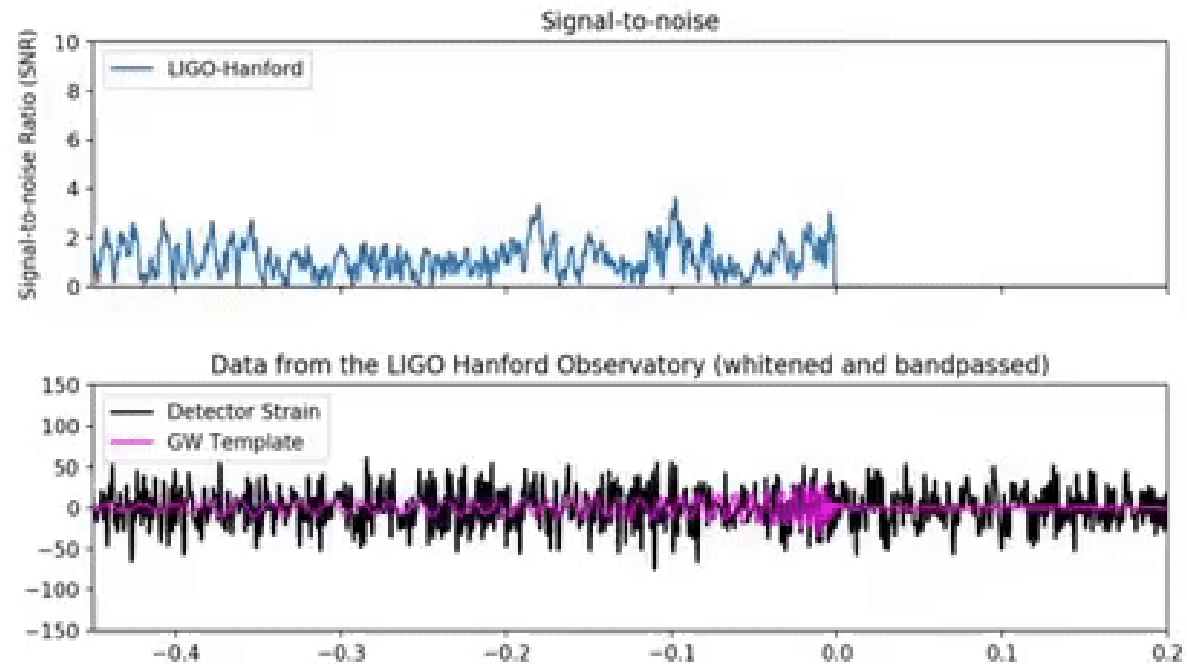
| Modeled searches | Unmodeled searches |
|------------------|--------------------|
| GSTLAL           | cWB                |
| MBTA             | oLIB               |
| PyCBC            |                    |
| SPIIR            |                    |

# Typical modeled search



GW150914

# Typical modeled search



GW151226

# Timeline

Time since gravitational-wave signal



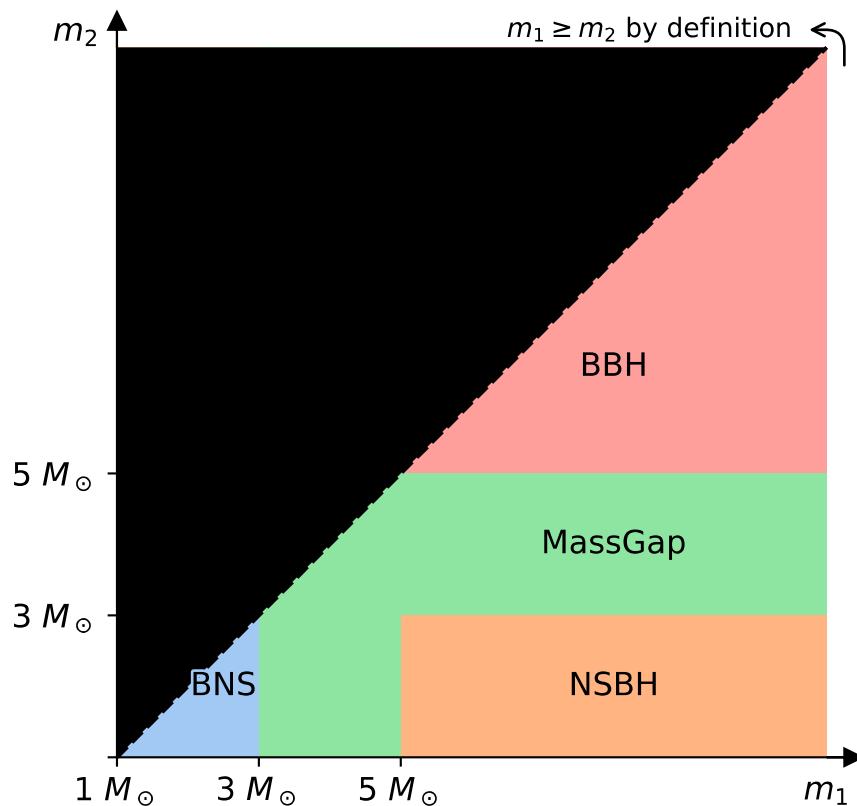
Several pipelines, each can make several detections

Based on a given criteria, one is chosen as being the **Preferred Event** (for the public alert)

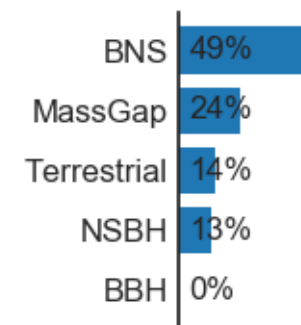
| Modeled searches | Unmodeled searches |
|------------------|--------------------|
| GSTLAL           | cWB                |
| MBTA             | oLIB               |
| PyCBC            |                    |
| SPIIR            |                    |

# Timeline

Time since gravitational-wave signal

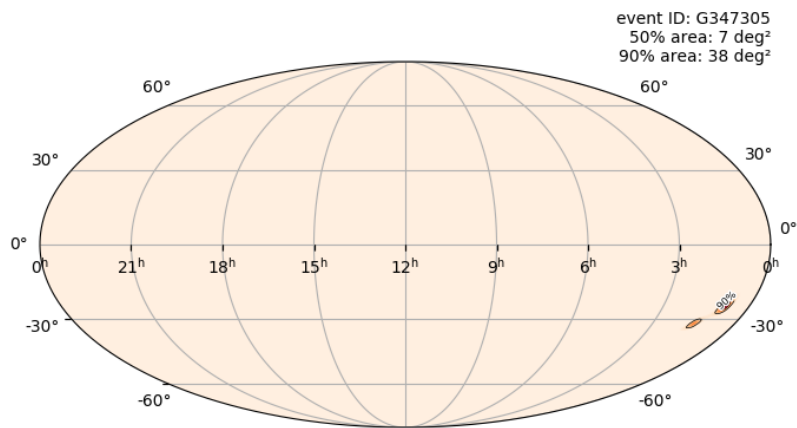


Example

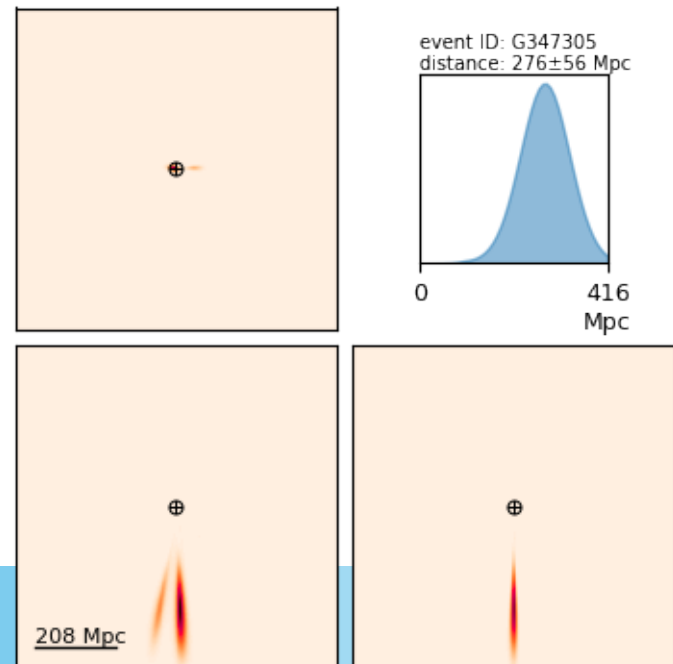


# Timeline

Time since gravitational-wave signal



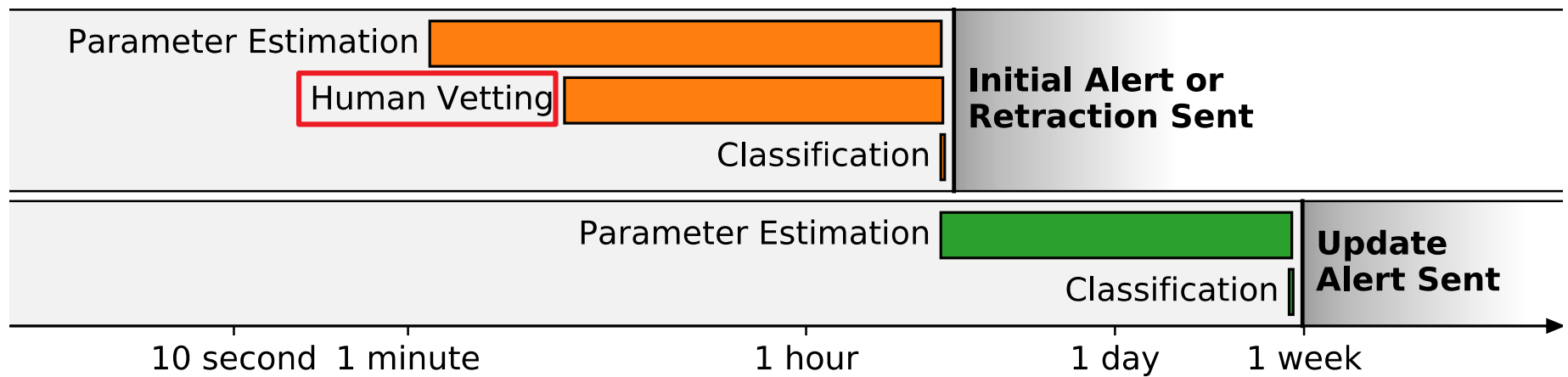
S190814bv (NSBH)





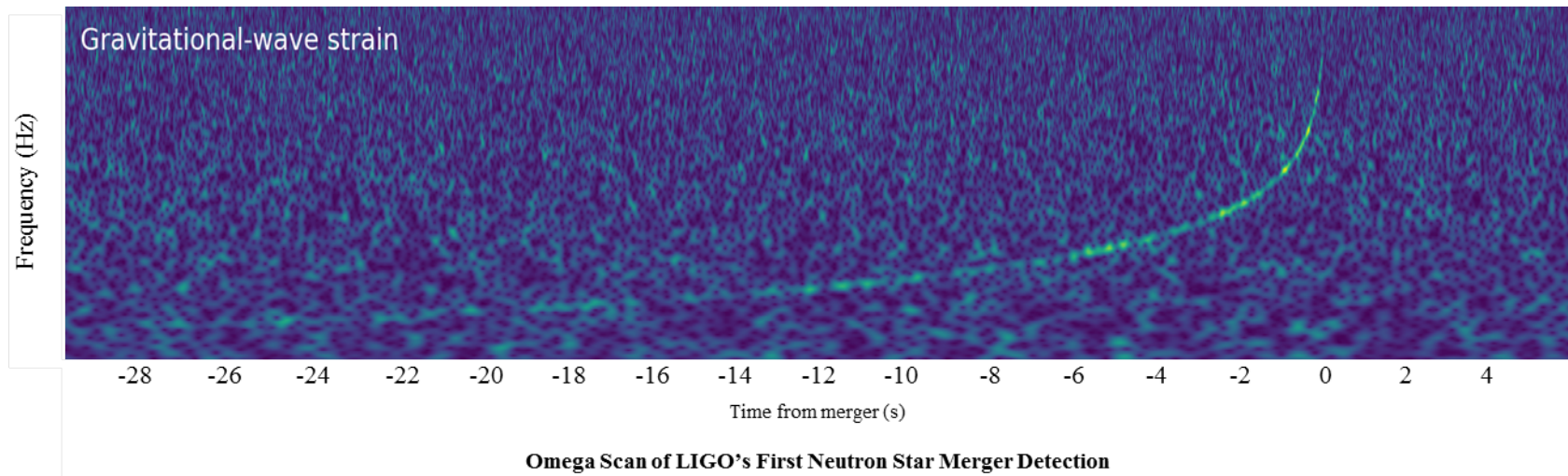
# Timeline

Time since gravitational-wave signal



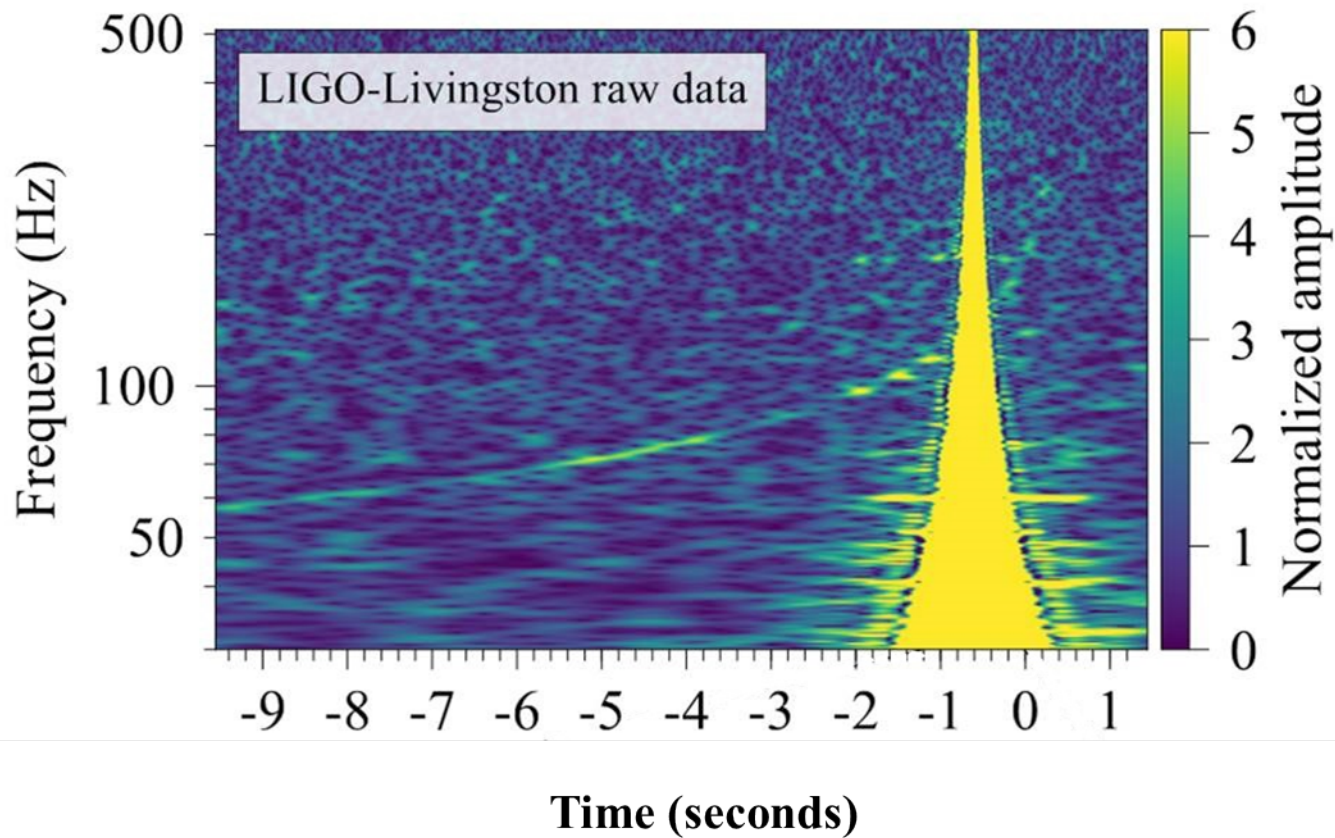
# Human vetting versus glitches

## Chirp for GW170817 (BNS)



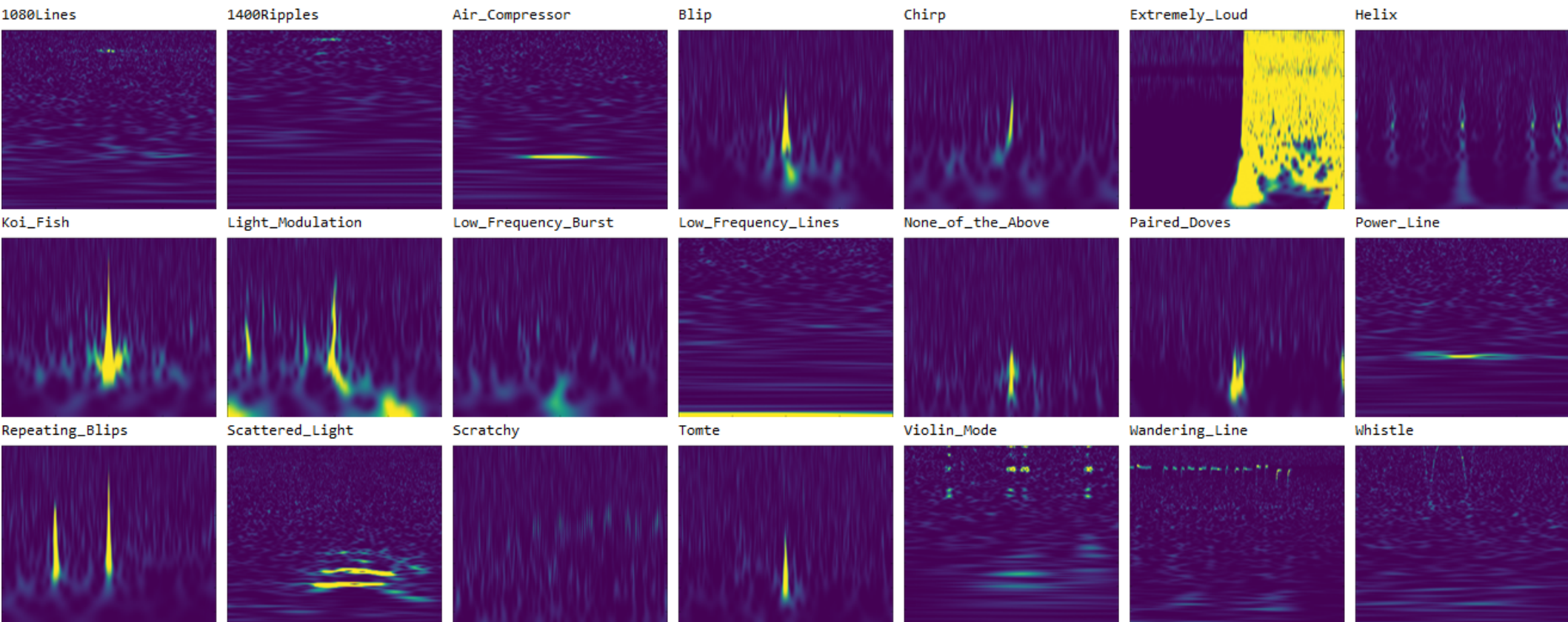
# Human vetting versus glitches

Chirp for GW170817 (BNS)

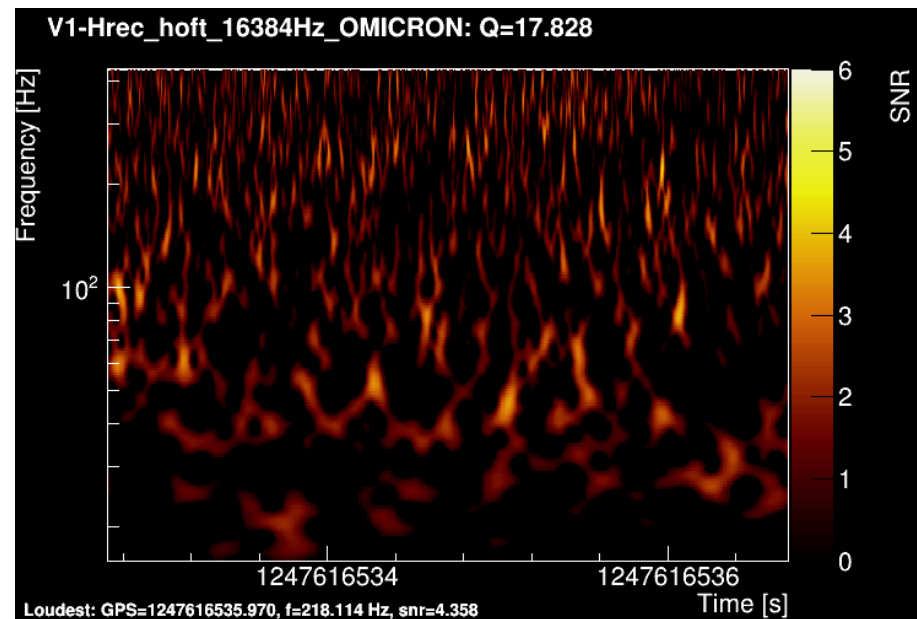
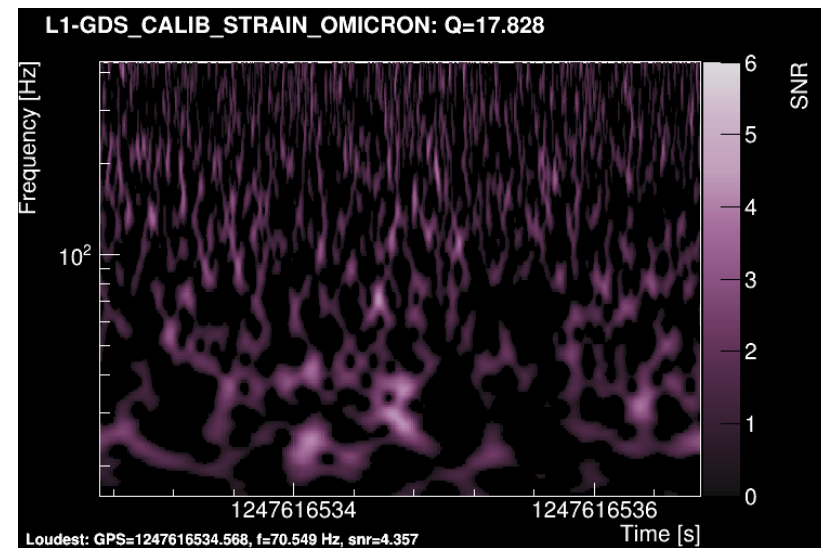
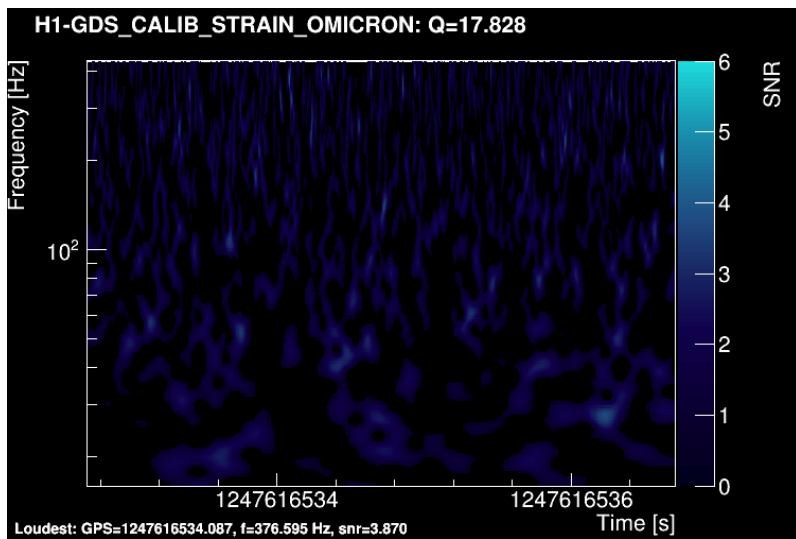


# Human vetting versus glitches

Anyone can help characterizing glitches at [Gravity Spy](#)  
(collaborative: provide training for IA)

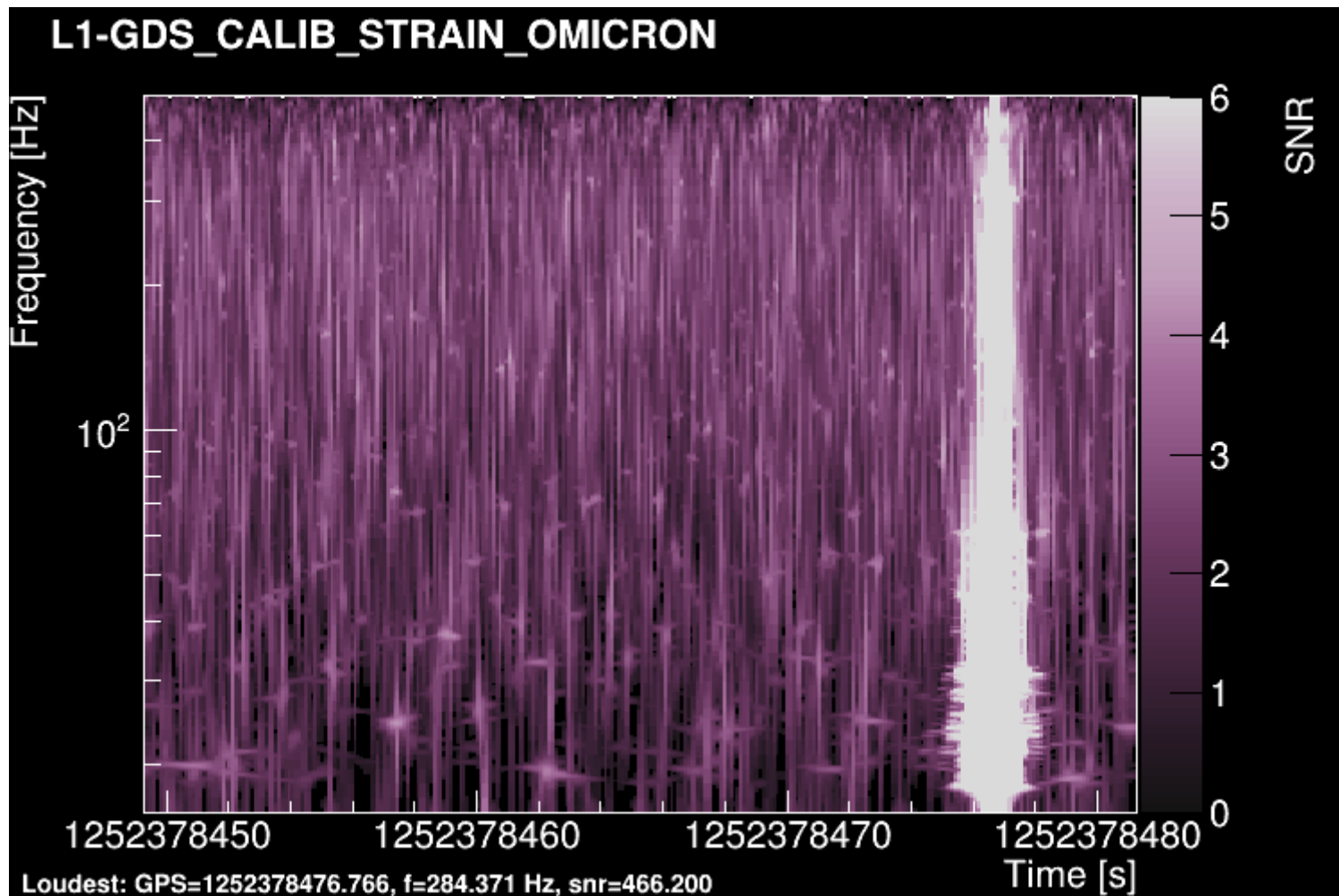


# Human vetting versus glitches



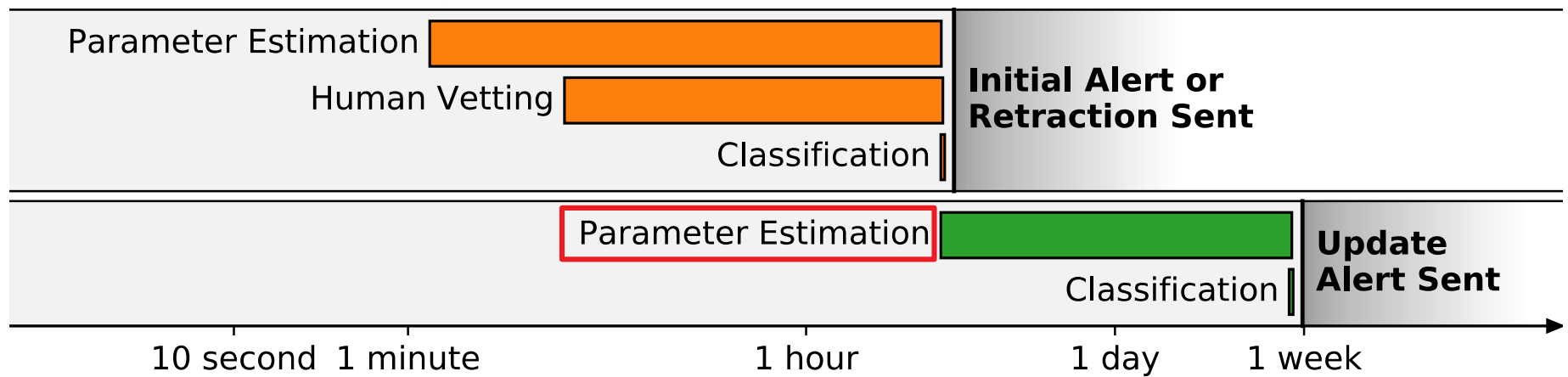


# Human vetting versus glitches



# Timeline

Time since gravitational-wave signal



# Parameter estimation: LALInference

## Bayesian parameter estimation

Markov Chain Monte Carlo (MCMC) or Nested sampling

### 15 parameters:

- 2 masses
- Luminosity distance
- Right ascension & declination
- Inclination angle
- Polarization angle
- Reference time
- Orbital phase at reference time
- 2 spin magnitudes
- Two angles for each spin

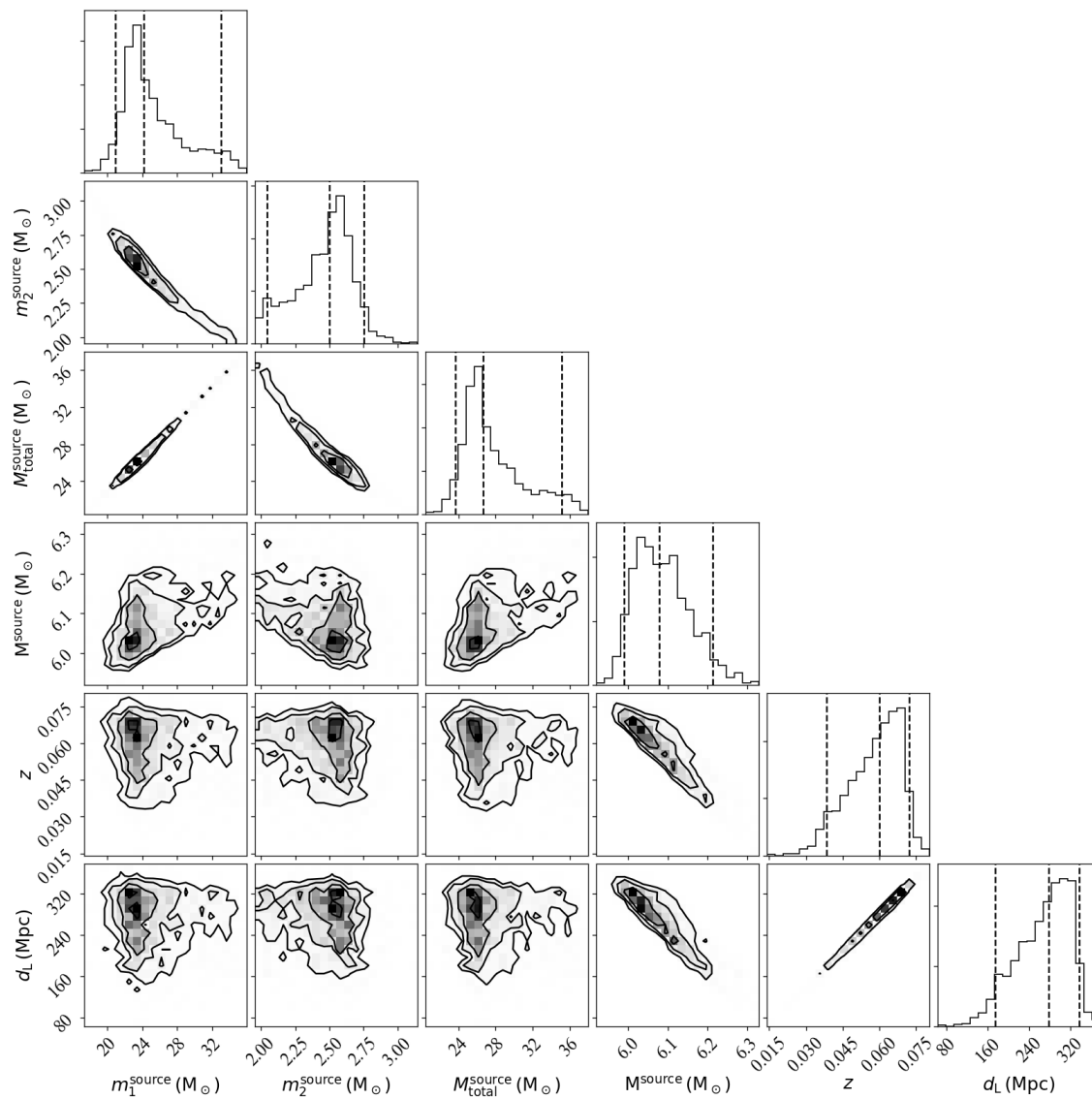
### Several models (under permanent improvement):

- Frequency domain
  - Post-Newtonian waveform
  - Phenomenological calibrated to numerical relativity
- Time domain
  - Post-Newtonian waveform
  - Effective One Body (EOB) → inspiral-merger-ringdown

With several EOS models for BNS and NSBH

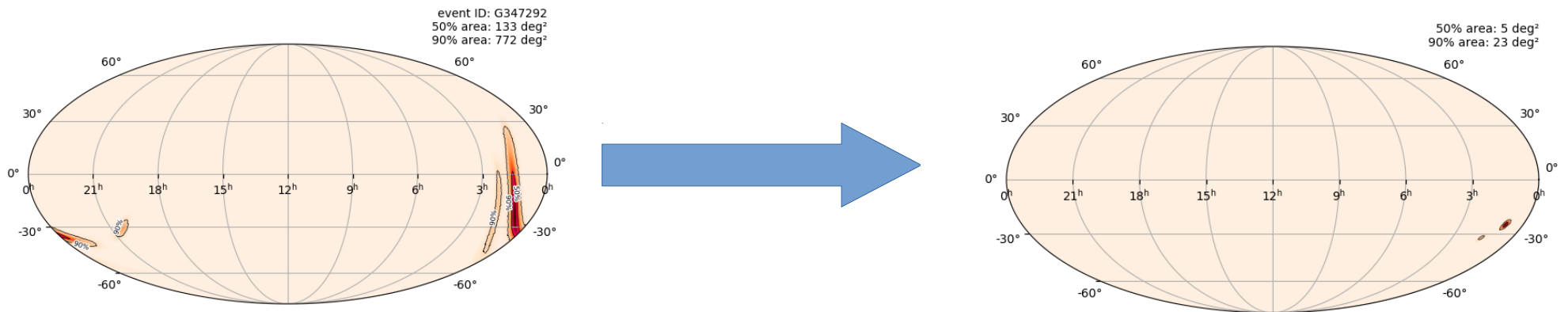


# Parameter estimation: LALInference



# Improved skymaps

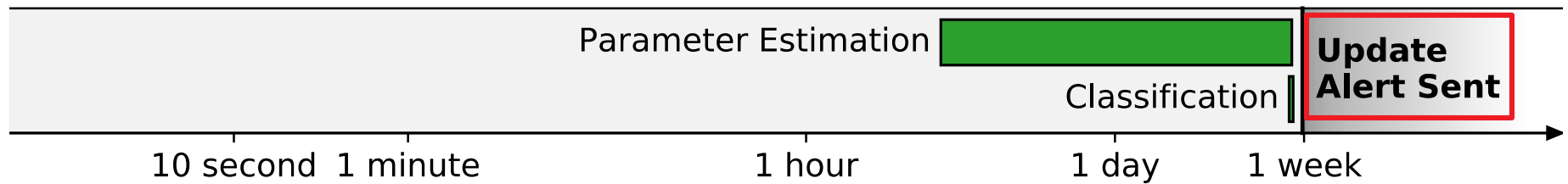
Sometimes significant improvement



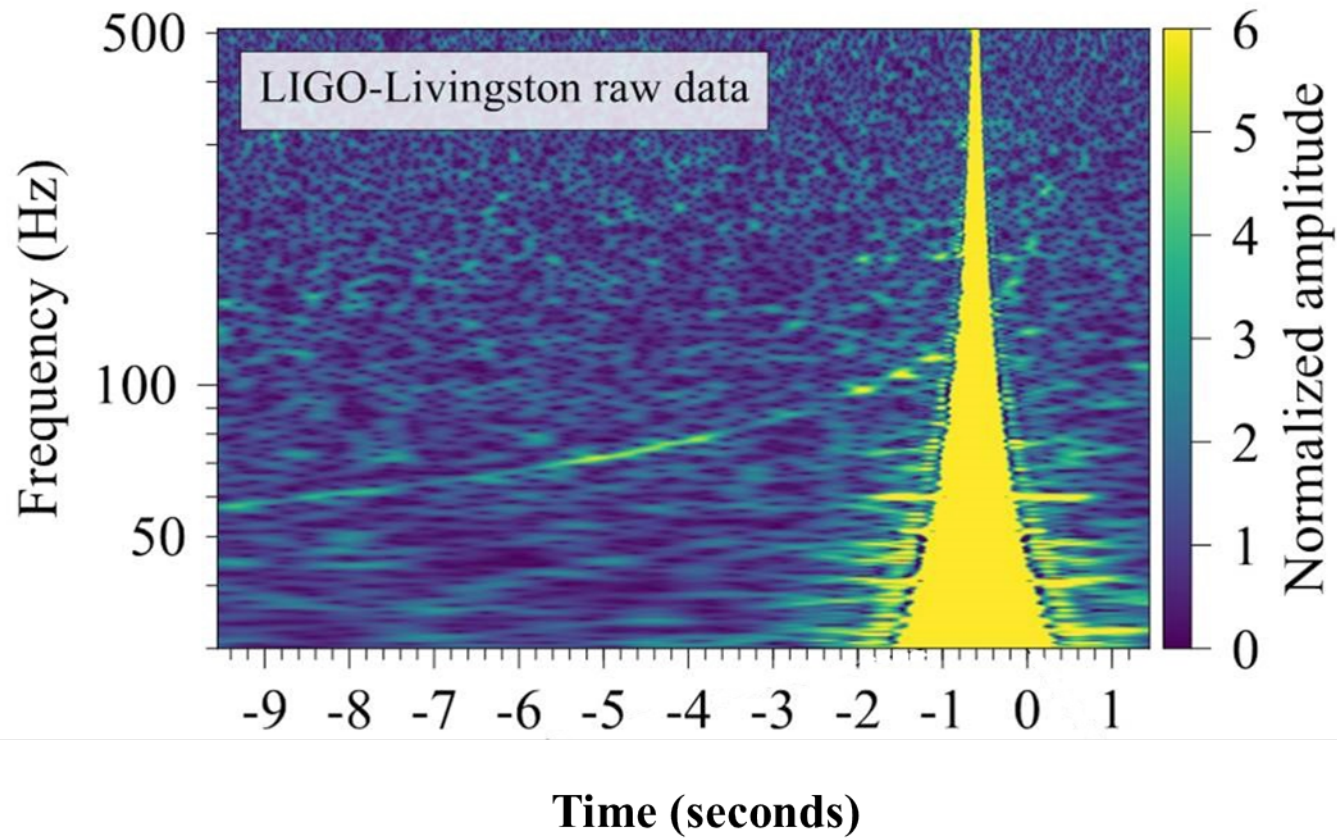
# Timeline

Time since gravitational-wave signal

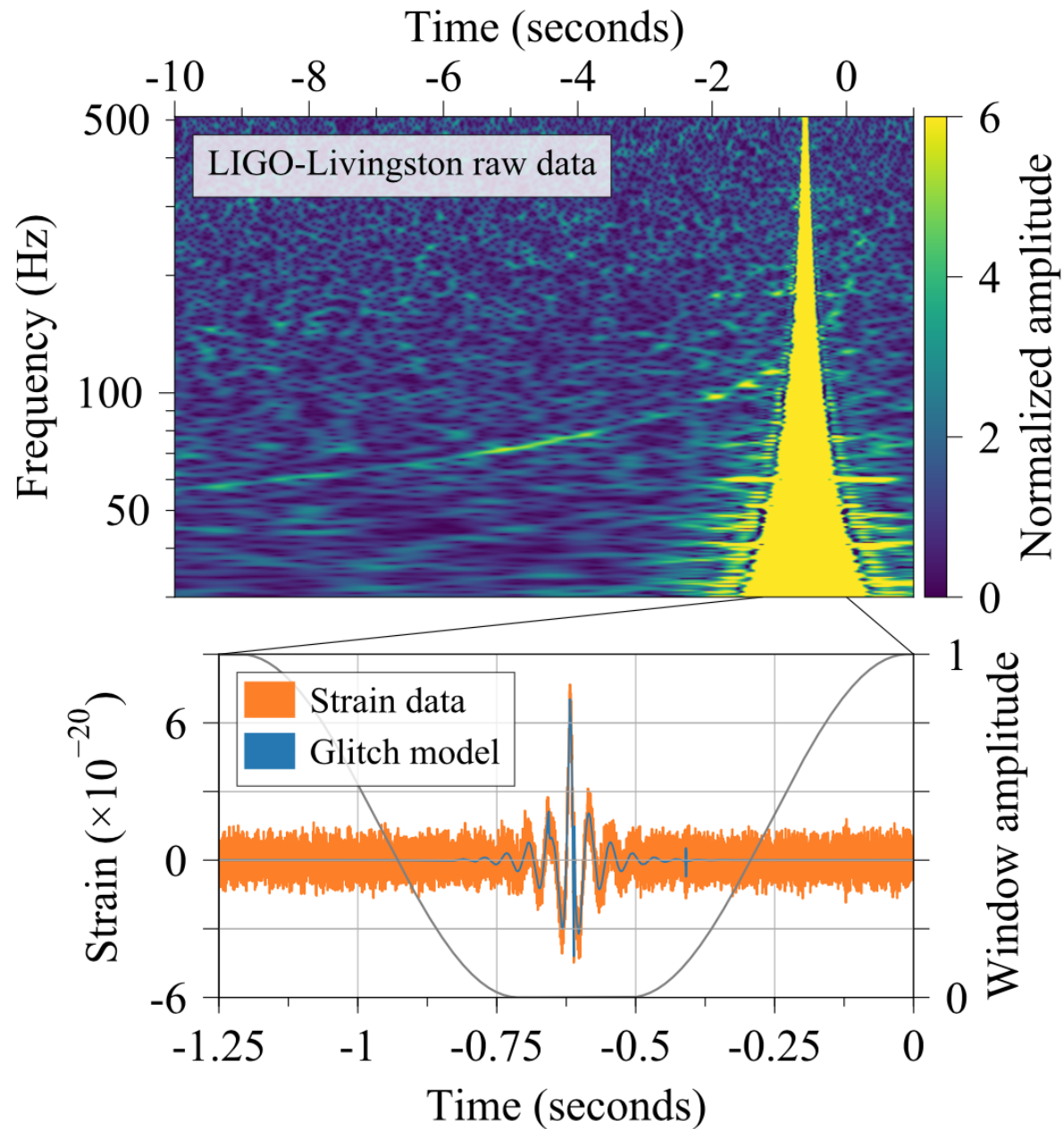
If better skymap and/or classification → update



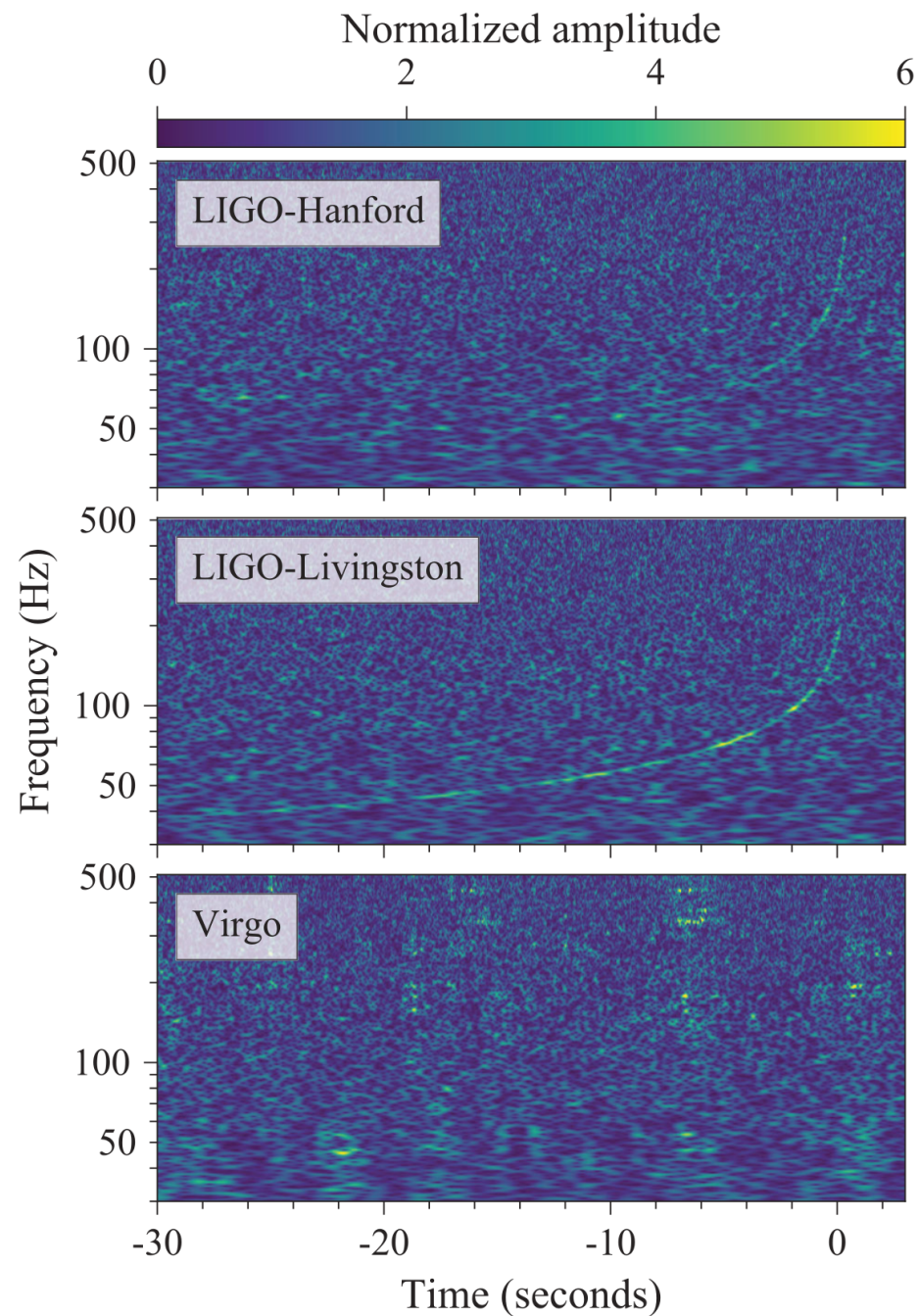
# Back to GW170817



# Back to GW170817

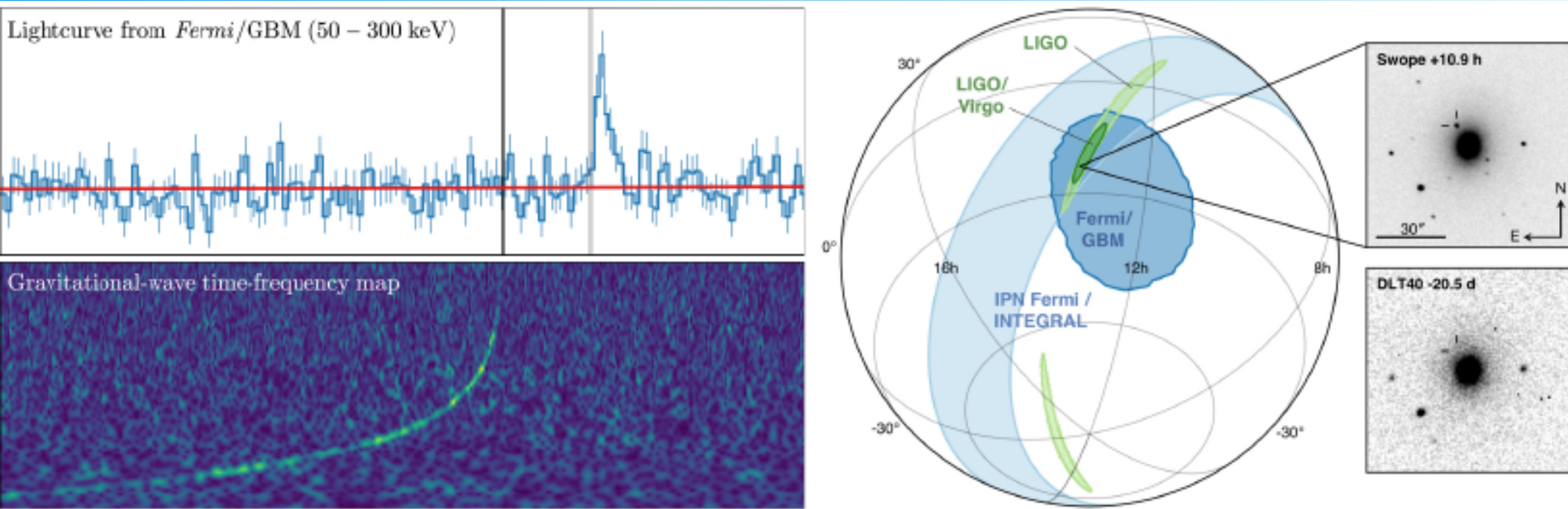


# Back to GW170817



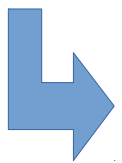


# END



All public O3 events can be found on GraceDB: <https://gracedb.ligo.org/superevents/public/O3/>

Tutorials and software open source on GWOSC: <https://www.gw-open-science.org/about/>  
(GWOSC: Gravitational Wave Open Science Center)



You can search for GWs signals in actual data from O1 and O2 with ~ one hour training!  
(and learn in preparation for O3)

# THANK YOU